CMS Barrel Muon Workshop

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First experience on DST analysis at T1/T2 centers Physics case study: $H \rightarrow WW \rightarrow 2\mu 2\nu$

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S. Lacaprara: First experience on DST analysis $H \to WW \to 2\mu 2\nu$ PRS/ μ session

Outline

Introduction,

- Analysis models on DC04 data,
- ► Future plan for data access,
- ► Physics case: $H \rightarrow WW \rightarrow 2\mu 2\nu$,
- Conclusion

Introduction

- ► During last two months Data Challenge '04
- Goal: demonstrate the capability of the CMS computing system to deal with events at 25 Hz in a distributed way
- Need of many millions of MC data, to be used also for physics study
- Large effort to setup best possible generation, detector simulation and response, reconstruction (from summer 2003)
- The MC production (aka PreDC04) was long, painful and not yet ended (only a fraction of event required by PRS have been fully processed)

- All MC data available at CERN (T0)
- ► Process MC data at 25 Hz to produce DST (at T0)
- Ship DST (plus streams) to "interested" T1 (eg CNAF, PIC, RAL, FNAL)
- ► Further move some DST from T1 to T2 (CNAF→LNL)
- ► Lots of files to be moved around (O(10⁶)): hard bookkeeping, bottleneck, etc, ...
- Last step is accessing data for analysis
- Still missing/on going/painful !!
- Not yet clear the "CMS analysis model": ie how to access data by the final user
- "Fake" or "real" analysis use-cases

Fake Analysis

- ► Base idea: synchronous with data arrival
- As soon a new run fully arrive at Tn, run a "standard" executable on it
- Implementation is very tricky: not meant (at least at the beginning) to be real (name says much...)
- Every run is not attached to MetaData, so they are fully independent. Not possible to run on a full dataset (or on a fraction of it), only run per run. "Deep Winter Mode" (aka Ice Age mode)
- Need real expert help, usage of several COBRA tools, parsing results, setup proper .orcarc, etc...
- Set up for PRS b/\(\tau\) and \(\mu\): running on several datasets

Pros:

- it works! (huge effort needed)
- can run a defined analysis (== executable) as soon as new data arrive
- May be a use case if you just need to increase statistics for your analysis

Cons:

- ★ definitively not for generic user!
- ★ Very complex!
- * not suited for developing analysis
- not easy (or even possible) to run many times on full dataset, only *incremental* analysis
- Large overhead for smallness of analyzed set of data

Real Analysis

- Base idea: access the data as you (generic, average user) always did
- No need of special expertise: just what the ORCA tutorial teach you
- Access dataset as such (not just a bunch of runs, but a well defined collection of them)
- Access full available data (or a fraction of them) as many time as you need (cpu power only limit!)
- Not necessarly real time analysis, but not too much delay either!
- Guarantee access to all users to all datasets (at least in principle)

Needs

◊ Data!

- Attached MetaData (collection of runs)
- Up-to-date catalog (where the files are)
- User access to farm where data are
- Needed information available to users: what data are available (dataset/owner/contents) and where they are (catalogs)

Ideas

- ★ Tn should work with catalog, not user!
- As soon as data are available, Tn prepare a catalog with attached MetaData to be used by user
- ★ User put this catalog and dataset/owner into
 - .orcarc
- ⋆ Run his/her analysis at Tn

Problems and status

- Many problems to have attached MetaData!! Large latency, very late availability (only very few datasets)
- Create a catalog for O(10⁴) files (fraction of a dataset) not easy: long and fragile
- Data integrity is crucial! Found problems on transfer agent: not easy to check data integrity

IT WORKS

for a fraction of a dataset mu03_DY2mu, managed to have all stuff setup and to run on ~ 80 kev (then crash for missing file... data available $\sim 10^6$, $\frac{1}{3}$ of full dataset)

Off Site Access

- what about non LNL user (namely all but me?)
- two possibilities:
 - Give to each user login to LNL. Done for DAQ TDR, painful for user support!
 - Use Grid tool to gain access to LNL
- ► Grid access works!
- Need to know how to setup CMS software, where catalog are, etc, as usual!
- Need some training on GRID (edg) tools (such as submission, output retrieval, job status query, etc). Not really hard

Future

- Tn produces a catalog for each dataset it has
 A
- Publish the catalog to RLS (grid filesystem)
- Our of the second seco
- Can have standard tool to perform such task: work on-going in Padova (N. Smirnov Grape)
- Use result of your query to decide where to run (in case a dataset is available in different Tn) use grid data
 different Tn) use grid data

discovery capability

- Use physical location of local catalog (of chosen site) in user .orcarc
- Run the jobs and get back the results

Status

- First dataset of DC04 data is becoming available at LNL for true analysis
- Only fraction of the dataset
- Many problems with data integrity: not possible (yet) to run on all available data, only small fraction
- Not clear when other dataset will become available (bottleneck is preparation of attached MetaData)
- Access to data via grid service is possible
- In principle anyone who can edg-job-submit, can look at the data
- Need to learn some grid tools, which are not as good as one would wish...
- Some ideas for short term improvement of grid access, with increasing exploiting of grid functionality

Physics case study: $H \rightarrow WW \rightarrow 2\mu 2\nu$

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Foreword

- All results presented based on DC04 data, running on DST only at LNL
- Run mainly to demonstrate data accessibility, "physics" is a by product (so far!)
- Very few result, due to lack of data
- Only two background dataset $t\bar{t} \rightarrow 2\mu$, $Z/\gamma^* \rightarrow 2\mu$ available for running a *standard* executable, no signal available!
- Just a demonstration that data are (veeeeery slooowly) becoming available...

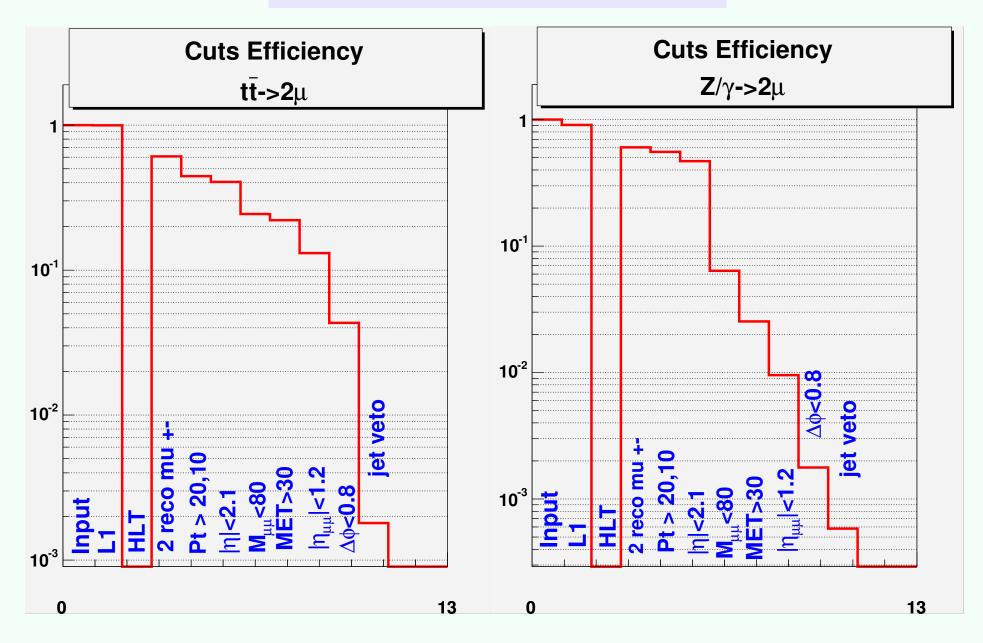
$H \rightarrow WW \rightarrow 2\mu 2\nu$

- Signal: 2 isolated μ, MET, no jets, μμ topology due to WW spin correlation (Δφ), no peak in invariant mass plot, or similar
- **Background:** all topologies with 2μ in final state
- $\blacktriangleright WW, WZ, Z/\gamma^{\star}, b\overline{b}, t\overline{t}, \text{ single top}, \ldots \rightarrow 2\mu + X$
- Set of cuts to reduce backgound
- Need large background reduction and reliable background estimation: counting experiment!
- Today: only two dataset partially available for analysis, and only few days ago! No signal!
- $\blacktriangleright t \bar{t} \rightarrow 2 \mu \sim 430$ kev: fake analysis at LNL
- ► $Z/\gamma^{\star} \rightarrow 2\mu$ 82 kev: true analysis with grid access to LNL from Padova

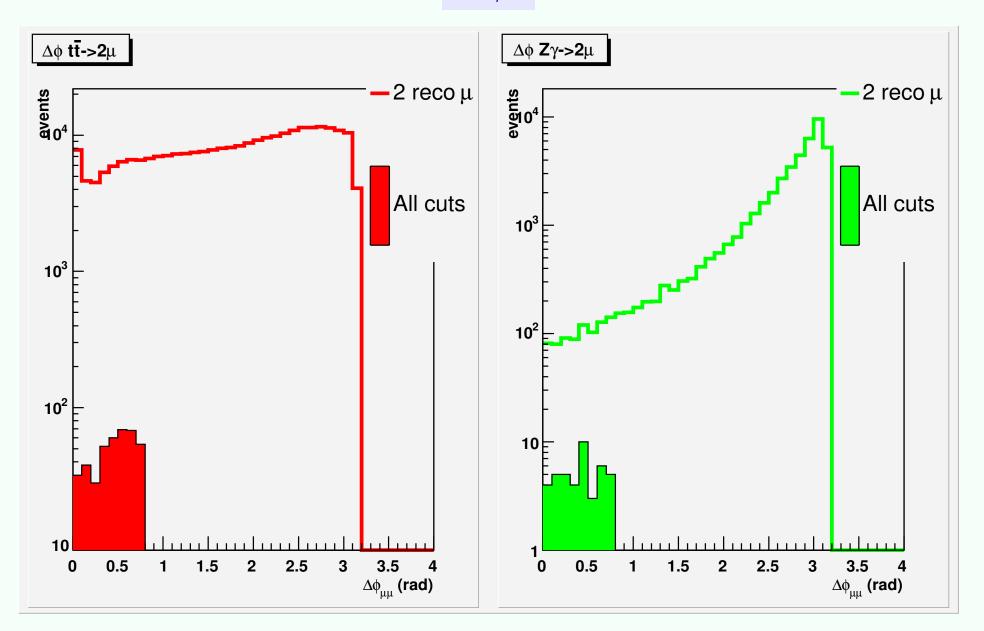
Cut Reminder

- ► L1 and HLT selection for single or di-muons,
- exactly 2 opposite sign muons reconstructed,
- ▶ $p_t^1 > 20 \; {\rm GeV}/c, \, p_t^2 > 10 \; {\rm GeV}/c,$
- ► $|\eta_{\mu}^{1,2}| < 2.1$ central muons,
- ► $M_{inv}^{\mu\mu} < 80 \text{ GeV}/c^2 \text{ Z peak cut},$
- ► MET > 20 GeV neutrinos,
- ► $|\eta_{\mu\mu}| < 1.1$ central events,
- ► $\Delta \phi < 0.8$, $H \rightarrow WW$ spin correlation,
- ► No jet $p_t > 30$ in $|\eta| < 2.4$ (top events)

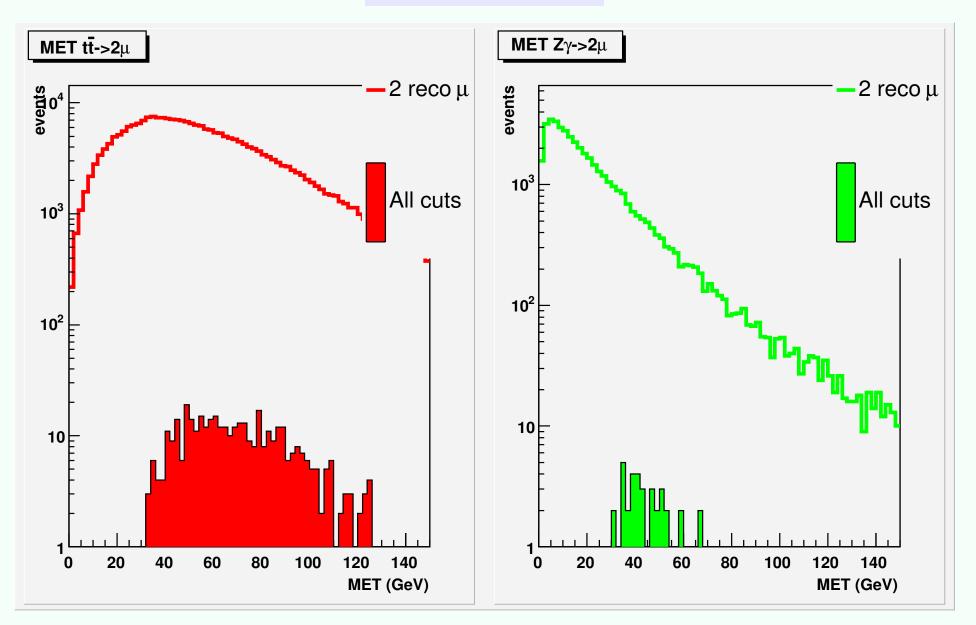
Background Rejection



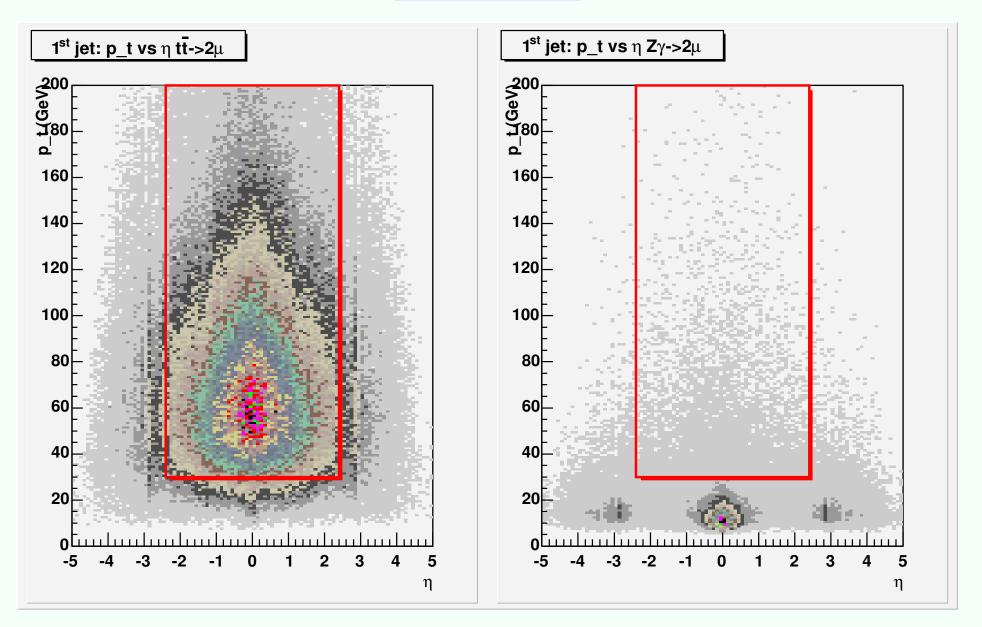
 $\Delta \phi$



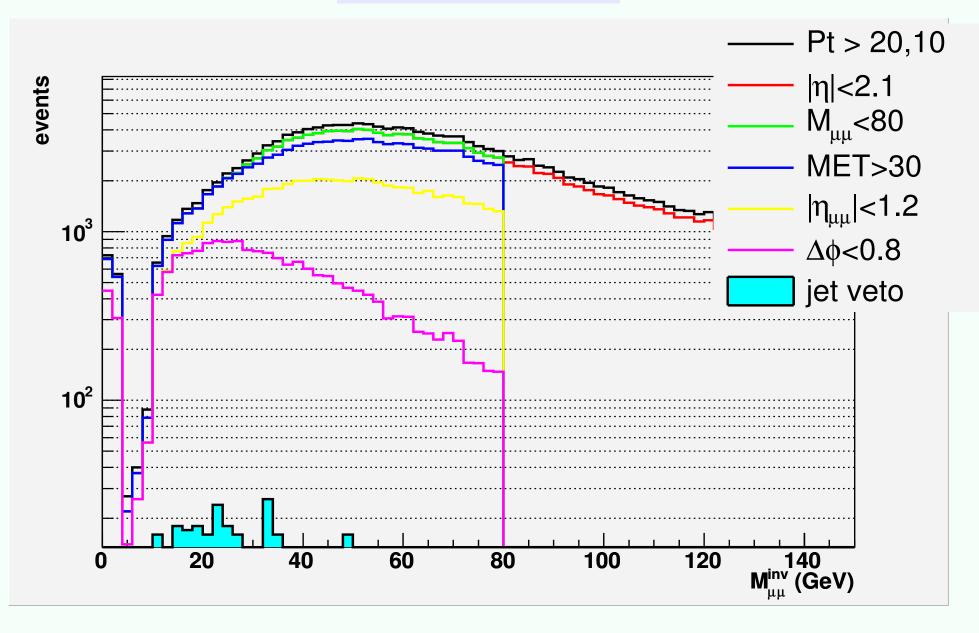
Missing E_t



Jet Veto

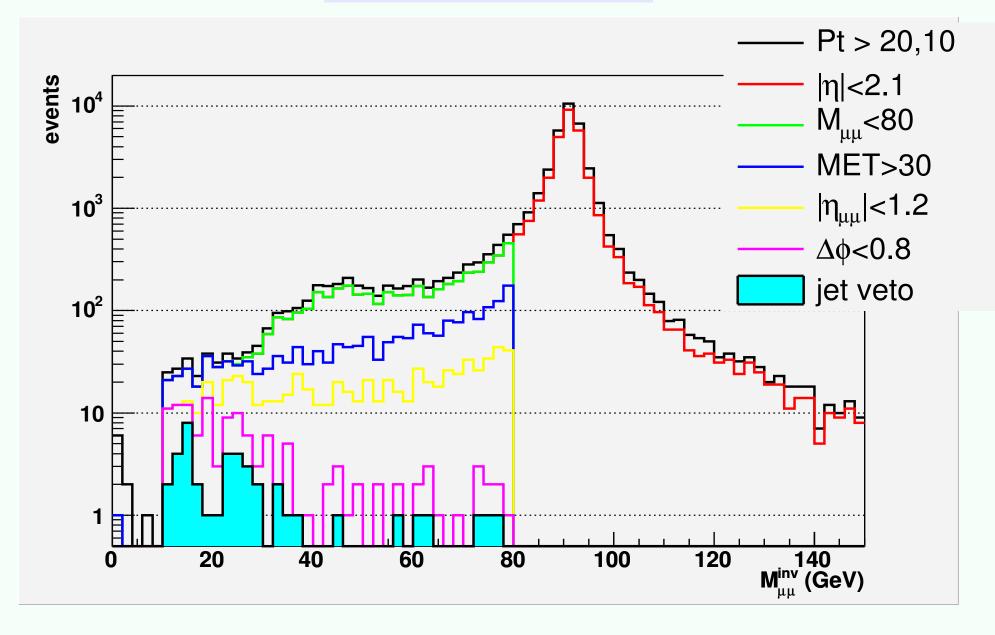


 $M^{inv}_{\mu\mu} t\bar{t} \rightarrow 2\mu$



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 $M^{inv}_{\mu\mu} Z\gamma^{\star} \rightarrow 2\mu$



Residual Rate

Very preliminary

Dataset	$t\bar{t} \rightarrow 2\mu$	$Z/\gamma^{\star} \to 2\mu$	
#ev "analyzed" (kev)	430	82	
full dataset (kev)	600	3000	
fraction	70%	3%	
Residual Rate (10^{-6} Hz)	32	17	
#ev left	770	33	
σ_{stat}	4%	17%	
σ_{stat} full DS	3%	3%	

* $H \rightarrow WW \rightarrow 2\mu 2\nu$ event rate (after cuts) from pre DC04 analysis (M.Zanetti thesis)

$M_{Higgs}~{ m GeV}/c^2$		140		
Residual Rate (10^{-6} Hz)	1.0	5.6	14	4.6

Conclusions

- Not a real analysis, yet: too few time to look at too few data!
- Test accessibility of DST data from DC04 shows first success
- "Results" presented here obtained running ORCA on DST (just reading, very fast!) from DC04 production, submitting jobs over the grid
- Need more effort to allow full PRS community to be able to do the same: timescale can be short
- Need much more data as soon as possible!!!
- A lot already fully processed, need to be made accessible

- Have a lot to learn how to access and use efficiently DST (feedback)
- Same for all different reconstruction algorithms (muons, jets, MET, vertexes, tracks, etc): problems, performances, tuning...
- ► Very first look at two backgrounds for $H \rightarrow WW \rightarrow 2\mu 2\nu$: need lot of work but promising!!
- Statistics (will be) available from DC04 enough for small statistical error (real work will be to estimate systematics)
- Is it finally time to start analysis??